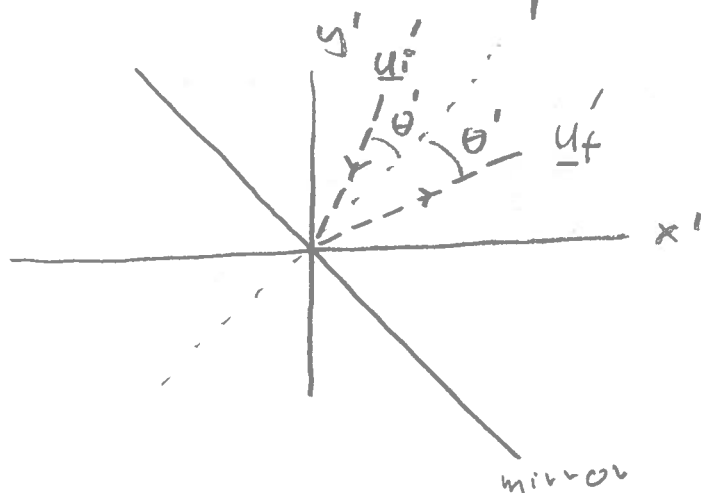


Solutions to Assignment 1

PHYS 514
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(1)

Q1 In S' the setup is



The initial velocity is \underline{u}'_i and the final one is \underline{u}'_f . Clearly

$$\tan\left(\frac{\pi}{4} - \theta'\right) = \frac{u'_{f,y}}{u'_{f,x}} \quad \text{and} \quad \tan\left(\frac{\pi}{4} + \theta'\right) = \frac{u'_{i,y}}{u'_{i,x}}$$

where $\frac{\pi}{4}$ is the angle between the normal to the mirror and the x' -axis.

It's easy to see that $u'_{i,y} = -u'_{f,x}$ and

$$u'_{i,x} = -u'_{f,y}$$

Let's now go to the system S which moves with velocity $-v$ in the y' -direction as seen from S' .

The usual velocity addition formulas give (2)
that

$$u_{i,y} = \frac{u'_{i,y} + v}{1 + v u'_{i,y}} \quad , \quad u_{i,x} = \frac{u'_{i,x}}{\gamma(1 + v u'_{i,y})}$$

$$u_{t,y} = \frac{u'_{t,y} + v}{1 + v u'_{t,y}} = \frac{-u'_{i,x} + v}{1 - v u'_{i,x}}$$

and

$$u_{t,x} = \frac{u'_{t,x}}{\gamma(1 + v u'_{t,y})} = \frac{-u'_{i,y}}{\gamma(1 - v u'_{i,x})}$$

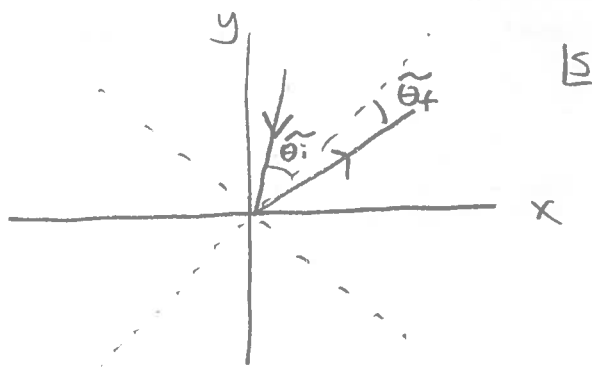
where $\gamma = \frac{1}{\sqrt{1-v^2}}$ and $c=1$.

Then it's easy to calculate the angles between the velocity vectors and the $x=y$ line (the normal to the $x=-y$ plane). They are

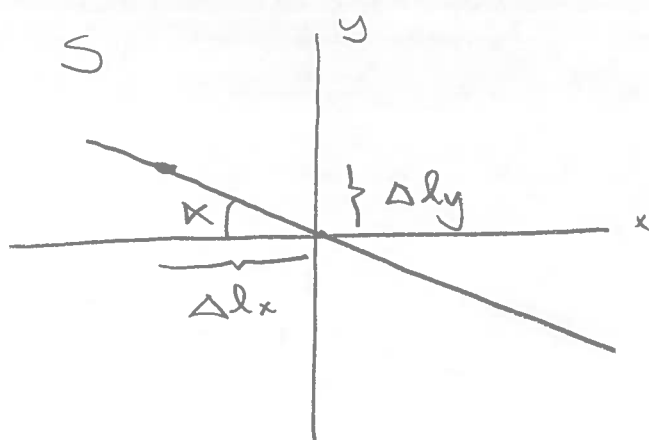
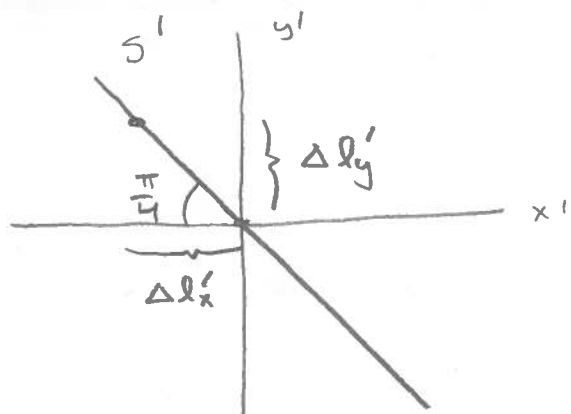
$$\tan\left(\frac{\pi}{4} + \tilde{\theta}_i\right) = \frac{u_{i,y}}{u_{i,x}} = \frac{u'_{i,y} + v}{\sqrt{1-v^2} u'_{i,x}}$$

and

$$\tan\left(\frac{\pi}{4} - \tilde{\theta}_t\right) = \frac{u_{t,y}}{u_{t,x}} = \frac{u'_{i,x} - v}{\sqrt{1-v^2} u'_{i,y}}$$



We must also take into account length contraction of the mirror. (3)



The usual formulas give that

$$\Delta l_x = \Delta l'_x \quad \text{and} \quad \Delta l_y = \sqrt{1-v^2} \Delta l'_y$$

where $\Delta l'_x = \Delta l'_y$. Thus

$$\alpha = \arctan \frac{\Delta l_y}{\Delta l_x} = \arctan \sqrt{1-v^2}$$

It's then easy to see that the incoming and outgoing angles relative to the normal of the mirror in S are

$$\begin{aligned} \theta_i &= \tilde{\theta}_i - \left(\frac{\pi}{4} - \alpha \right) = \arctan \left(\frac{u'_{iy} + v}{\sqrt{1-v^2} u'_{ix}} \right) - \frac{\pi}{4} - \frac{\pi}{4} + \alpha \\ &= \arctan \left(\frac{u'_{iy} + v}{\sqrt{1-v^2} u'_{ix}} \right) + \arctan \sqrt{1-v^2} - \frac{\pi}{2} \end{aligned}$$

and

$$\theta_f = \tilde{\theta}_f + \left(\frac{\pi}{4} - \alpha \right)$$

$$= -\arctan \left(\frac{u'_{ix} - v}{\sqrt{1-v^2} u'_{iy}} \right) + \frac{\pi}{2} - \arctan \sqrt{1-v^2}$$

This can be rewritten using expressions like

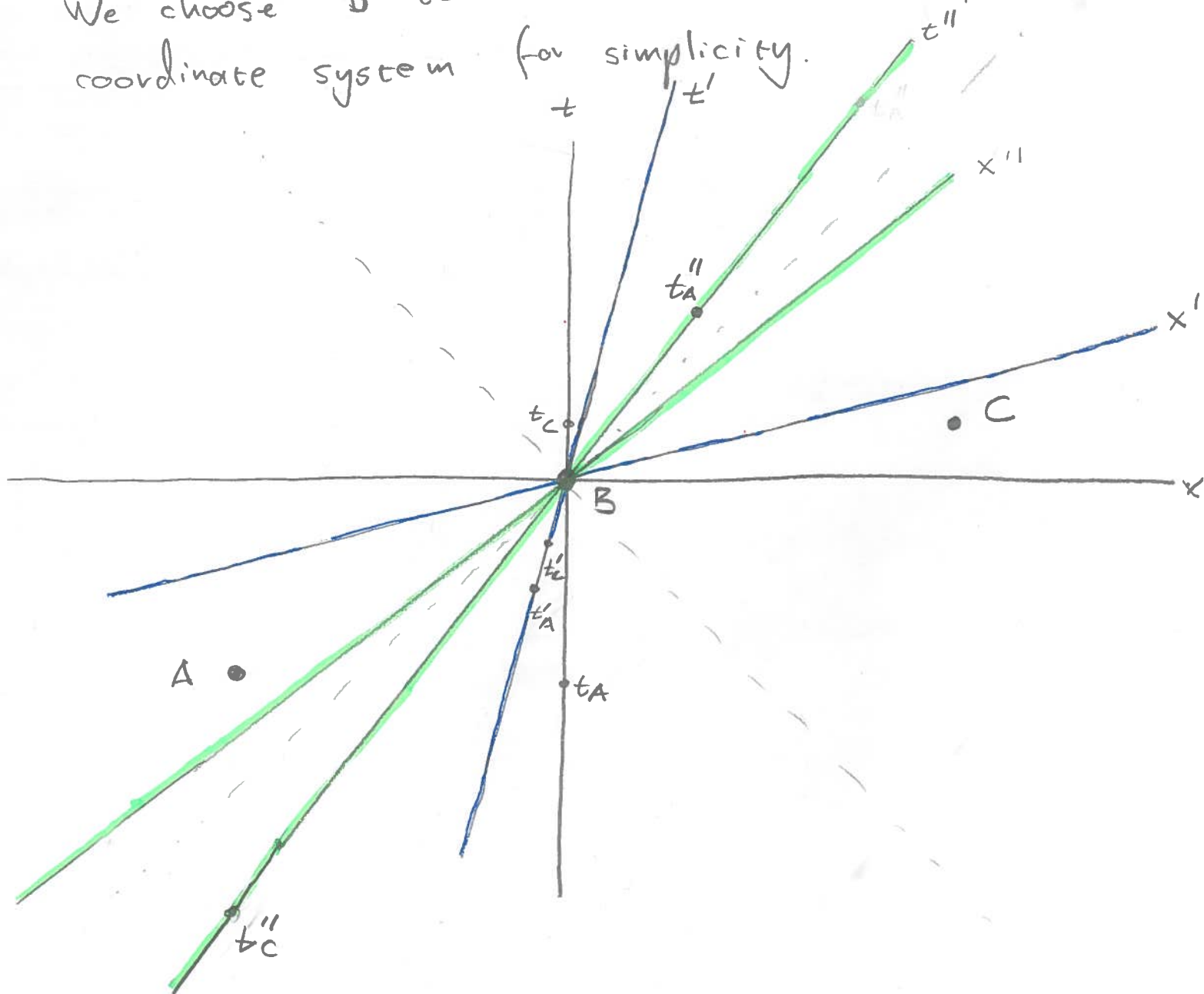
$$\arctan \frac{1}{x} = \frac{\pi}{2} - \arctan x,$$

$$\arctan x + \arctan y = \arctan \left(\frac{x+y}{1-xy} \right)$$

but that doesn't offer much simplification.

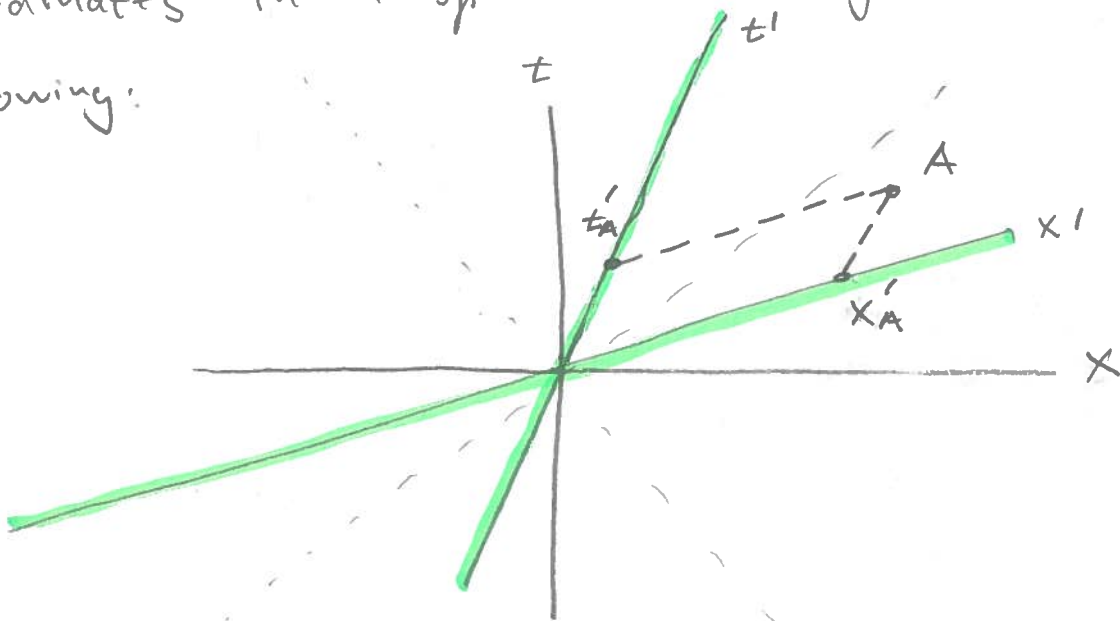
Q2: Yes, this is possible for some combination of three events.

We choose B to be at the center of the coordinate system for simplicity.



We see that in S the order is ABC , in S' (5) it is ACB and in S'' it is CBA .

For clarification: The correct way to see coordinates in a space-time diagram is the following:



Q3: The muon mass is $m = 0.106 \text{ GeV}$, its energy is $E = 1000 \text{ GeV}$ and the ^{rest-frame} lifetime is $\tau = 2.19 \times 10^{-6} \text{ s}$. Since $E = \gamma m$ the time dilation formula tells us that the lifetime in the lab frame is

$$t = \gamma \tau = \frac{E \tau}{m} = \underline{\underline{0.021 \text{ s}}}$$

It travels distance ct in a storage ring of circumference πd where d is the diameter.

Thus it travels

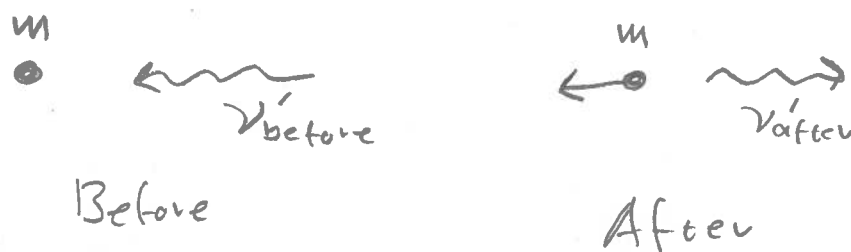
$$\frac{ct}{\pi d} \cdot 2\pi = \underline{\underline{1.3 \times 10^4 \text{ rad}}} \quad \text{before decaying.}$$

Q4: We can analyze this problem directly in the frame of reference S where $E \gg m \gg \nu$ by using momentum conservation.

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We can also take a shortcut by doing the analysis in the rest frame of the massive particle, S' , and then boosting.

In S'



The usual formula for Compton scattering says that

$$\nu'_{\text{after}} = \frac{\nu'_{\text{before}}}{1 + \frac{\nu'_{\text{before}}}{m} (1 - \cos \theta)}$$

where θ is the angle between the initial and final direction of motion of the photon. We will clearly maximize the energy transfer if

$$\theta = \pi \quad \text{so} \quad \nu'_{\text{after}} = \frac{\nu'_{\text{before}}}{1 + 2 \frac{\nu'_{\text{before}}}{m}}$$

(If $\theta = 0$ there is no scattering which is clearly not the case we're interested in.)

Boosting with velocity v we get that

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$$\gamma'_{\text{after}} = \gamma \gamma_{\text{after}} (1-v)$$

and

$$\gamma'_{\text{before}} = \gamma \gamma (1+v).$$

So that

$$\gamma \gamma_{\text{after}} (1-v) = \frac{\gamma \gamma (1+v)}{1 + 2 \frac{\gamma}{m} \gamma (1+v)}.$$

We have that the boost needed corresponds to

$$\gamma = \frac{E}{m} \quad \text{and since } E \gg m \text{ we also have}$$

$$v \approx 1.$$

Thus

$$\begin{aligned} \gamma_{\text{after}} &= \gamma \frac{\frac{\gamma (1+v)^2}{1-v^2}}{1 + 2 \frac{\gamma}{m} (1+v) \gamma} \\ &\approx \gamma \frac{4 E^2 / m^2}{1 + 4 \frac{E}{m^2} \gamma} = \gamma \frac{4 E^2}{m^2 + 4 E \gamma}. \end{aligned}$$

Thus the energy transfer is

$$\gamma_{\text{after}} - \gamma \approx \gamma_{\text{after}} \approx \frac{4 E^2 \gamma}{m^2 + 4 E \gamma}$$

(b) Here $E = 10^{20} \text{ eV}$ and

$$M = 9.4 \times 10^8 \text{ eV.}$$

There are many ways to understand what is meant by a 3K photon. Let's just use the

estimate
$$h\nu = k_B T$$

so that
$$\nu = \frac{k_B T}{2\pi} = \frac{8.6 \times 10^{-5} \frac{\text{eV}}{\text{K}} \cdot 3\text{K}}{2\pi} = 0.41 \times 10^{-4} \text{ e}$$

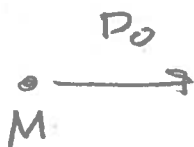
where we use that $\frac{h}{2\pi} = 1$.

Then the energy transfer is

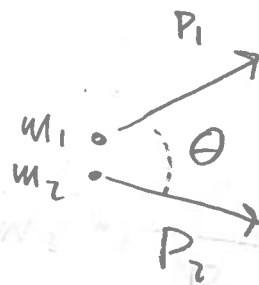
$$\frac{4E^2\nu}{M^2 + 4E\nu} \approx \underline{\underline{1.8 \times 10^{14} \text{ eV}}}$$

Q5:

Before



After



Here P_0, P_1, P_2 are four-vectors and

$P_1 = (E_1, \underline{p}_1)$, $P_2 = (E_2, \underline{p}_2)$. We have that

$$P_0^2 = -M^2, \quad P_1^2 = -m_1^2, \quad P_2^2 = -m_2^2.$$

Energy-momentum conservation gives $P_0 = P_1 + P_2$

$$\text{so } P_0^2 = (P_1 + P_2)^2 = P_1^2 + 2P_1 \cdot P_2 + P_2^2,$$

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$$\text{i.e. } -M^2 = -m_1^2 - m_2^2 - 2E_1 E_2 + 2p_1 \cdot p_2,$$

$$\text{i.e. } M^2 = \underline{m_1^2 + m_2^2 + 2E_1 E_2 - 2p_1 \cdot p_2 \cos\theta}$$

Q6: (a) The non-relativistic action is

$$S_{\text{nr}} = \int \frac{1}{2} m v^2 dt.$$

$$\text{Thus } 0 = \delta S_{\text{nr}} = \int m v^i \delta v^i dt = \int m v^i \frac{d x^i}{dt} dt$$

$$\stackrel{\substack{\nearrow \\ \text{integration} \\ \text{by parts}}}{=} - \int m \frac{d v^i}{dt} \delta x^i dt$$

Clearly the equation of motion is $\dot{v}^i = \frac{d v^i}{dt} = 0$.

(b) The relativistic action is $S = m \int dz = \int m \frac{dt}{\gamma}$

where τ is the eigentime and t is the lab time.

Thus $S = \int m \sqrt{1-v^2} dt$. In the non-relativistic

limit $\sqrt{1-v^2} \approx 1 - \frac{1}{2} v^2$ so

$S = \int m dt - S_{\text{nr}}$, i.e. we get the non-relativistic action up to a constant (and an irrelevant minus sign).

In general

$$0 = \delta S = \int m \frac{-v^i}{(1-v^2)^{1/2}} \delta v^i dt$$

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$$= - \int m \gamma v^i \frac{d\delta x^i}{dt} dt = \int m \frac{d}{dt} (\gamma v^i) dt \cdot \delta x^i$$

Thus the equation of motion is simply

$$\frac{d}{dt} (\underline{p}) = 0 \quad \text{where} \quad \underline{p} = \gamma m \underline{v}$$

is the momentum.

Since $\gamma \rightarrow 1$ in the non-relativistic limit we get the result of (a).